# ON THE T-MATRIX FOR WATER-WAVE SCATTERING PROBLEMS

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A rigid cylinder of infinite length is floating in the free surface of deep water. The cylinder is held fixed and a given time-harmonic wave of small amplitude is incident upon it. The corresponding linear two-dimensional boundary-value problem for a velocity potential  $\phi$  is treated using the null-field method, and an expression for the T-matrix is obtained. (The T-matrix connects the diffraction potential away from the cylinder to the given incident potential.) Fundamental properties of the T-matrix are derived from considerations of energy and reciprocity. For regular wavetrains incident from the right or from the left, there are well-known relations between the corresponding reflection and transmission coefficients; these relations are recovered by specialising the equations satisfied by the T-matrix. Two extensions to water of constant finite depth are described: one uses multipole potentials whilst the other uses Havelock wavemaker functions; this second approach also leads to a new method for treating the problem of waves in a semi-infinite channel with an end-wall of arbitrary shape.

#### 1. Introduction

A rigid cylinder of infinite length is floating in the free surface of deep water. The cylinder is held fixed and a given time-harmonic wave of small amplitude is incident upon it. This wave is scattered and induces hydrodynamic forces on the cylinder. To model this situation, we make the usual assumptions of classical hydrodynamics [1]: we assume that the water is incompressible and inviscid, that the motion is irrotational, and that the wave crests are parallel to the generators of the cylinder. Consequently, we obtain the following linear, two-dimensional boundary-value problem for a velocity potential  $Re\{\phi(P) e^{-i\omega t}\}$ .

Scattering problem S

Find a function  $\phi(P)$  such that  $\phi$  satisfies Laplace's equation,

$$\left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2}\right) \phi = 0 \quad \text{in } D, \tag{1.1}$$

the free-surface condition

$$K\phi + \frac{\partial \phi}{\partial y} = 0 \quad \text{on } F \tag{1.2}$$

and the boundary condition

$$\frac{\partial \phi}{\partial n} = 0 \quad \text{on } \partial D, \tag{1.3}$$

 $\phi - \phi_I \equiv \phi_D$  satisfies a radiation condition, and  $|\text{grad }\phi| \to 0$  as  $y \to \infty$ . (The exact kinematic and dynamic boundary conditions on the moving free surface have been linearised and combined to give (1.2).)

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Here, D, F and  $\partial D$  denote the fluid domain, the mean free surface and the wetted surface of the cylinder, respectively. x and y are Cartesian coordinates, chosen so that the y-axis is vertical (y increasing with depth) with F occupying a portion of the line y=0; the remaining portion of this line is denoted by  $F_-$  and is assumed to contain the origin O.  $D_-$  denotes the interior of the cylinder, i.e. the region with boundary  $\partial D \cup F_-$ . Capital letters P, Q denote points of D, lower-case letters P, P0 denote points of P2, and P3, and P4 denote points of P5. See Fig. 1.

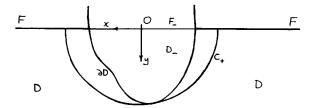


Fig. 1. The floating cylinder.

 $K = \omega^2/g$ , where  $\omega$  is the radian frequency and g is the acceleration due to gravity, is a given positive number.  $\phi$ ,  $\phi_1$  and  $\phi_D$  are called the total, incident and diffraction potentials, respectively.  $\phi_1$  is given, satisfies (1.1) everywhere in y > 0 (except possibly at a finite number of isolated points in D) and (1.2) everywhere on y = 0.

John [2] has proved that, under certain geometrical restrictions on  $\partial D$ , S is uniquely solvable for all values of K. We shall henceforth assume that these restrictions are met.

One method for solving S is the null-field method [3, 4]. This requires the solution of an infinite set of moment-like equations for  $\phi(q)$  (these are called the *null-field equations*); an integral representation then gives  $\phi(P)$  for all  $P \in D$ . Let  $C_+(C_-)$  be the escribed (inscribed) semicircle to  $\partial D$ , centered on O. Then, for P outside  $C_+$ , we can formally reduce the null-field equations and the integral representation to

$$c_m = \sum_{n=1}^{\infty} T_{mn} d_n, \quad m = 1, 2, \ldots$$

where  $d_n$  are the (known) coefficients in a certain expansion of  $\phi_1(P_-)$  for  $P_-$  inside  $C_-$  and  $c_m$  are the coefficients in a certain expansion of  $\phi_D(P)$  for P outside  $C_+$  (see Section 3).  $T_{mn}$  is usually known as the *T-matrix*. Such a matrix was first introduced in the theory of quantum scattering [5, Section 7.2.2] and was later used by Waterman [6] to treat electromagnetic scattering problems; it is now used widely in classical scattering theory [7].

In this paper, we shall derive some fundamental properties of the T-matrix for S, namely

$$T_{mn} = T_{nm} \tag{1.4}$$

and

$$\frac{2}{\pi}\operatorname{Im}(T_{mn}) + T_{1m}T_{1n}^* + T_{2m}T_{2n}^* = 0, \tag{1.5}$$

where the asterisk denotes the complex conjugate. These results are obtained (in Section 4) from reciprocity and energy considerations for S. In fact, it is well known that if these considerations are valid for *any* given scattering process, then similar results will obtain; see, e.g. [8].

The results (1.4) and (1.5) can be used to provide independent checks on numerical calculations. It should also be possible to use (1.4) and (1.5) as constraints in a numerical method for finding approximations to T; this approach has been used successfully by Waterman [8] and Werby and Green [9].

In Section 5, we consider specific incident waves, namely regular waves incident from  $x = +\infty$  or from  $x = -\infty$ ; let  $r_{\pm}$  and  $t_{\pm}$  denote the corresponding complex reflection and transmission coefficients. We show that the known relationships between these coefficients [10] can be recovered by specialising (1.4) and (1.5).

In the last two sections, we consider the more complicated situation of finite depth. We describe two approaches: the first is a natural generalisation of the infinite-depth approach, but leads to a complicated analogue of (1.5); the second approach uses Havelock wavemaker functions—besides yielding a relation similar to (1.5), it also leads to a novel method for treating problems involving a semi-infinite channel with an end-wall of arbitrary shape. Possible extensions to three dimensions are described.

### 2. Reciprocity and energy relations

Let  $\phi_1$  and  $\phi_2$  be two scattering potentials, corresponding to two incident potentials,  $\phi_{11}$  and  $\phi_{12}$ , i.e.

$$\phi_i = \phi_{1i} + \phi_{Di}, \quad i = 1, 2$$

where  $\phi_i$  solves S. Choose two distinct points,  $P_1$  and  $P_2$ , in D, and set

$$\phi_{1i}(P) = G(P; P_i), \quad i = 1, 2$$
 (2.1)

where G(P; Q) is the potential at P due to a simple wave source at Q in the absence of the body (G is given explicitly by (3.1), below). By Green's theorem, we have

$$2\pi\{\phi_2(P_1)-\phi_1(P_2)\}=[\phi_1,\phi_2]-[\phi_1,\phi_2]_C$$

where

$$[\phi_1, \phi_2]_C = \int_C \left(\phi_1 \frac{\partial \phi_2}{\partial n} - \phi_2 \frac{\partial \phi_1}{\partial n}\right) ds, \tag{2.2}$$

 $[\phi_1, \phi_2] \equiv [\phi_1, \phi_2]_{\partial D}$ , C is any simple path in D connecting the positive x-axis to the negative x-axis,  $\partial/\partial n$  denotes normal differentiation on C in the outward direction (i.e. away from the body), and it is assumed that  $P_1$  and  $P_2$  lie in the fluid between C and  $\partial D$ . By (1.3),  $[\phi_1, \phi_2] = 0$ . Also, since  $\phi_1$  and  $\phi_2$  both satisfy the radiation condition (i.e. not merely  $\phi_{D1}$  and  $\phi_{D2}$ ),  $[\phi_1, \phi_2]_C \to 0$  as the path C recedes to infinity. Hence

$$\phi_2(P_1) = \phi_1(P_2). \tag{2.3}$$

Since G is symmetric, we have

$$\phi_{12}(P_1) = G(P_1; P_2) = G(P_2; P_1) = \phi_{11}(P_2),$$

whence (2.3) reduces to

$$\phi_{\rm D2}(P_1) = \phi_{\rm D1}(P_2).$$
 (2.4)

This reciprocity relation states that the diffraction potential at  $P_1$  due to a wave source at  $P_2$  is the same as the diffraction potential at  $P_2$  due to a wave source at  $P_1$ . Reciprocity relations are also well known in other branches of scattering theory; see, e.g., [11, Section 1.32].

Note that the derivation of (2.4) is independent of the water depth, which need not be constant (the symmetry of G can be proved, using Green's theorem, without knowing G explicitly). Note also that

 $\phi_1(P)$  is not known; if it were, we could apply Green's theorem to  $\phi_D$  and  $\phi_1$  and obtain

$$\phi_{\mathrm{D}}(P_{\mathrm{I}}) = \frac{-1}{2\pi} \int_{\partial D} \phi_{\mathrm{I}}(q) \frac{\partial}{\partial n_{q}} \phi_{\mathrm{I}}(q) \, \mathrm{d}s_{q},$$

which is a formula for  $\phi_D$  ( $\phi_1$  is known as the exact Green function, or Neumann function, for S; see, e.g. [12, p. 38]).

Let us now consider the average flux of energy through a control contour C, namely

$$E = \frac{1}{2}\rho\omega \text{ Im } \int_{C} \phi^* \frac{\partial \phi}{\partial n} ds = -\frac{1}{4}i\rho\omega [\phi^*, \phi]_{C}.$$

If  $\phi_1$  is regular inside C, Green's theorem gives

$$E = -\frac{1}{4}i\rho\omega[\phi^*, \phi] = 0,$$

by (1.3), i.e. the average flux of energy through C is zero. Writing  $\phi = \phi_I + \phi_D$  and noting that  $[\phi_1^*, \phi_I] = 0$  (because  $\phi_I$  is regular in  $D_-$ ), we obtain

$$0 = \frac{1}{2}\rho\omega \operatorname{Im}[\phi_1^*, \phi_D] + E_D \tag{2.5a}$$

where

$$E_{\rm D} = -\frac{1}{4}i\rho\omega[\phi_{\rm D}^*, \phi_{\rm D}] \tag{2.5b}$$

is the average flux of energy through C due to the diffracted waves only. We shall refer to (2.5) as the energy relation.

The energy relation is usually stated for the special case in which  $\phi_1$  corresponds to a train of regular surface waves. Moreover, it is usually expressed in terms of Kochin's *H*-function; for details, see, e.g., [1, p. 249] or [10]. Similar relations have been obtained in other fields, where they are known as 'forward-scattering' theorems, or 'cross-section' theorems, or 'optical' theorems; see, e.g., [11, Sections 8.6, 8.24] or [13-16].

### 3. The null-field equations and the T-matrix

The potential of a simple wave source is [17]

$$G(P,Q) \equiv G(x,y;\xi,\eta) = \frac{1}{2}\log\frac{(x-\xi)^2 + (y-\eta)^2}{(x-\xi)^2 + (y+\eta)^2} - 2\int_0^\infty e^{-k(y+\eta)}\cos k(x-\xi)\frac{dk}{k-K},$$
 (3.1)

where the path of integration passes below the pole of the integrand at k = K. We apply Green's theorem in D to  $\phi_D$  and G, and in  $D_-$  to  $\phi_I$  and G, and then add the resulting equations to give

$$2\pi\phi_{D}(P) = -\int_{\partial D} \phi(q) \frac{\partial}{\partial n_{q}} G(P, q) \, ds_{q}$$
 (3.2)

and

$$2\pi\phi_{\mathbf{I}}(P_{-}) = \int_{\partial D} \phi(q) \frac{\partial}{\partial n_{q}} G(P_{-}, q) \, \mathrm{d}s_{q}. \tag{3.3}$$

(3.2) is an integral representation for  $\phi_D$ , valid for all  $P \in D$ . (3.3) is valid for all  $P_- \in D_-$ .

G has a bilinear expansion [17],

$$G(P,Q) = \sum_{m=1}^{\infty} \alpha_m(P) \Phi_m(Q)$$
(3.4)

for  $r_P < r_Q$ , where the functions  $\alpha_m$  and  $\Phi_m$  are defined in Appendix A; each is harmonic and satisfies the free-surface condition (1.2);  $\alpha_m$  are regular, whilst the *multipole potentials*  $\Phi_m$  are singular at O and satisfy the radiation condition. Henceforth, we shall simplify our notation and use a summation convention: sum over repeated suffices from 1 to  $\infty$ . Thus, (3.4) becomes

$$G(P, Q) = \alpha_m(P)\Phi_m(Q).$$

If we restrict P to lie outside  $C_+$  and  $P_-$  to lie inside  $C_-$ , we can substitute (3.4) into (3.2) and (3.3) giving

$$2\pi \phi_{\mathcal{D}}(P) = c_m \Phi_m(P) \tag{3.5}$$

and

$$2\pi\phi_{\mathbf{I}}(P_{-}) = d_{m}\alpha_{m}(P_{-}), \tag{3.6}$$

where

$$c_m = -\langle \phi, \alpha_m \rangle, \quad m = 1, 2, \dots, \tag{3.7}$$

$$d_m = \langle \phi, \Phi_m \rangle, \quad m = 1, 2, \dots \tag{3.8}$$

and

$$\langle f, g \rangle = \int_{\partial D} f(q) \frac{\partial g(q)}{\partial n_q} ds_q$$

The constants  $d_m$  (m = 1, 2, ...) are known; they are given in terms of  $\phi_1$  by

$$d_m = [\phi_1, \Phi_m]. \tag{3.9}$$

(3.8) are called the *null-field equations*; they are to be solved for  $\phi(q)$  and are known to be uniquely solvable at all frequencies [3, 4]. Once  $\phi(q)$  has been determined,  $\phi_D(P)$  is given by (3.2), or, when P is outside  $C_+$ , by (3.5) and (3.7).

To solve the null-field equations, we begin by choosing a basis for representing functions defined on  $\partial D$ ; let  $\{\phi_n(q)\}$ ,  $n = 1, 2, \ldots$ , be such a basis. Thus, we may write

$$\phi(q) = a_n \phi_n(q) \tag{3.10}$$

where  $a_n$ , n = 1, 2, ..., are unknown coefficients. Substituting (3.10) into (3.8) gives

$$Q_{mn}a_n = d_m, \quad m = 1, 2, \dots$$
 (3.11)

where

$$Q_{mn} = \langle \phi_n, \Phi_m \rangle. \tag{3.12}$$

Similarly, substituting (3.10) into (3.7) gives

$$c_m = -\hat{Q}_{mn}a_n, \quad m = 1, 2, \dots$$
 (3.13)

where

$$\hat{Q}_{mn} = \langle \phi_n, \alpha_m \rangle. \tag{3.14}$$

The system (3.11) is uniquely solvable, i.e.  $Q^{-1}$ , the inverse of the infinite matrix Q, exists. Thus, eliminating  $a_n$  between (3.11) and (3.13), we obtain

$$c_m = T_{mn}d_n, \quad m = 1, 2, \dots$$
 (3.15)

where

$$T_{mn} = -\hat{Q}_{ml}Q_{ln}^{-1} \tag{3.16}$$

is known as the *T-matrix*. Given T, we can determine the diffraction potential,  $\phi_D$ , outside  $C_+$  for any given incident potential,  $\phi_1$ , without computing the values of  $\phi$  on  $\partial D$ .

It is easy to see that the unique-solvability of S implies that T exists and is unique. This in turn implies that T is independent of the choice of basis  $\{\phi_n\}$ . However, this choice may be important in numerical calculations, when T must necessarily be truncated.

Numerical solutions of the null-field equations have been presented in [3, 4, 18]. In [3], simple choices for  $\{\phi_n\}$  are used. In [4], a different method is used; this method is convergent, but is only applicable when  $\phi_1$  corresponds to regular surface waves on deep water. Multiple-scattering problems are considered in [18].

In the next section, we shall derive two basic properties of the T-matrix. These are consequences of the reciprocity and energy relations given in Section 2, and do not depend on the *derivation* of the T-matrix given above; we merely require that the coefficients  $c_m$  and  $d_m$  occurring in the representations (3.5) and (3.6), respectively, are related through a matrix  $T_{mn}$  by (3.15).

# 4. Fundamental properties of the T-matrix

#### 4.1. Reciprocity relation

Let

$$2\pi\phi_{Di}(P) = c_m^{(i)}\Phi_m(P)$$
 and  $2\pi\phi_{Ii}(P_-) = d_m^{(i)}\alpha_m(P_-)$ 

for i = 1, 2, P outside  $C_+$  and  $P_-$  inside  $C_-$ . From (2.1) and (3.4), we obtain

$$d_m^{(i)} = 2\pi \Phi_m(P_i), \quad m = 1, 2, \ldots,$$

for i = 1, 2 and  $P_i \in D$ . Hence

$$2\pi\{\phi_{D2}(P_1) - \phi_{D1}(P_2)\} = c_m^{(2)}\Phi_m(P_1) - c_m^{(1)}\Phi_m(P_2) = \Phi_m(P_1)T_{mn}d_n^{(2)} - \Phi_m(P_2)T_{mn}d_n^{(1)}$$
$$= 2\pi\{\Phi_m(P_1)T_{mn}\Phi_n(P_2) - \Phi_m(P_2)T_{mn}\Phi_n(P_1)\}.$$

Thus, using the reciprocity relation (2.4), we obtain

$$0 = \Phi_m(P_1)\Phi_n(P_2)\{T_{mn} - T_{nm}\};$$

this identity must hold for all  $P_1$  and  $P_2$  outside  $C_+$ , whence

$$T_{mn}=T_{nm}, (4.1)$$

i.e. the T-matrix is symmetric.

# 4.2. Energy relation

From (3.5), we have

$$4\pi^{2}[\phi_{D}^{*},\phi_{D}] = 4\pi^{2}[\phi_{D}^{*},\phi_{D}]_{C_{A}} = c_{k}^{*}c_{i}[\Phi_{k}^{*},\Phi_{i}]_{C_{A}} = c_{k}^{*}c_{i}[\Phi_{k}^{*},\Phi_{i}].$$

Similarly, if  $\phi_1$  is regular inside  $C_+$ ,

$$4\pi^{2}[\phi_{1}^{*},\phi_{D}] = d_{k}^{*}c_{i}[\alpha_{k}^{*},\Phi_{i}],$$

where we have also used (3.6). Substituting into the energy relation (2.5), we obtain

$$0 = d_m^* d_n \{ T_{in} [\alpha_m^*, \Phi_i] - T_{im}^* [\alpha_n, \Phi_i^*] + T_{km}^* T_{in} [\Phi_k^*, \Phi_i] \}.$$

This identity must hold for any incident potential (that is regular inside  $C_{+}$ ) whence

$$T_{in}[\alpha_m^*, \Phi_i] - T_{im}^*[\alpha_n, \Phi_i^*] + T_{km}^* T_{in}[\Phi_k^*, \Phi_i] = 0.$$
(4.2)

From [17], we have

$$[\alpha_m^*, \Phi_j] = [\alpha_m, \Phi_j^*] = 2\pi \delta_{jm} \tag{4.3}$$

and

$$[\Phi_k^*, \Phi_i] = 2\pi^2 i \delta_{ik} (\delta_{i1} + \delta_{i2}) \quad \text{(no summation)}$$

where  $\delta_{ij}$  is the Kronecker delta. Thus, (4.2) becomes

$$T_{mn} - T_{nm}^* + \pi i (T_{1m}^* T_{1n} + T_{2m}^* T_{2n}) = 0$$

which, on using (4.1), reduces to

$$\frac{2}{\pi}\operatorname{Im}(T_{mn}) + T_{1m}T_{1n}^* + T_{2m}T_{2n}^* = 0. \tag{4.5}$$

Observe that only  $\Phi_1$  and  $\Phi_2$  can carry energy to infinity (these are the only two wavelike multipole potentials; see Appendix A); this accounts for the occurrence of the subscripts 1 and 2 in (4.5).

# 5. The Kreisel-Meyer relations

In this section, we shall consider particular incident waves, namely regular surface waves. We show that the well-known Kreisel-Meyer relations [10] can be obtained by specialising the fundamental relations satisfied by the *T*-matrix.

Consider a train of regular surface waves, propagating from  $x = +\infty$  towards the cylinder; the corresponding potential is

$$\phi_{I} = A_{+} e^{-Ky - iKx}$$

where  $A_+$  is a complex constant. This incident wave will be partially reflected and partially transmitted. We define  $r_+$  and  $t_+$  by

$$\phi(P) \sim \begin{cases} A_{+} e^{-Ky} (e^{-iKx} + r_{+} e^{iKx}) & \text{as } x \to +\infty, \\ A_{+} t_{+} e^{-Ky - iKx} & \text{as } x \to -\infty; \end{cases}$$

$$(5.1)$$

 $r_+$  and  $t_+$  are the (complex) reflection and transmission coefficients. Since  $\phi = \phi_1 + \phi_D$ , we have

$$\phi_{D}(P) \sim \begin{cases} A_{+}r_{+} e^{-Ky+iKx} & \text{as } x \to +\infty, \\ A_{+}(t_{+}-1) e^{-Ky-iKx} & \text{as } x \to -\infty. \end{cases}$$

$$(5.2)$$

 $\phi_{\rm D}$  is also given by (3.5); using the asymptotic behaviour of  $\Phi_m$  given in Appendix A, we find that

$$\phi_{D}(P) \sim \begin{cases} \frac{1}{2}(c_{1} + ic_{2}) e^{-Ky + iKx} & \text{as } x \to +\infty, \\ \frac{1}{2}(-c_{1} + ic_{2}) e^{-Ky - iKx} & \text{as } x \to -\infty. \end{cases}$$
 (5.3)

Comparing (5.2) and (5.3), we see that

$$c_1 = A_+(r_+ - t_+ + 1) (5.4)$$

and

$$c_2 = -iA_+(r_+ + t_+ - 1). (5.5)$$

Now, from (3.6) and (A.1), we have

$$d_1 = \mathrm{i} \pi A_+, \qquad d_2 = -\pi A_+$$

and  $d_n = 0$  for n > 2. Thus, from (3.15), we have

$$c_1 = \pi A_+ (i T_{11} - T_{12})$$
 and  $c_2 = \pi A_+ (i T_{21} - T_{22})$ ,

whence (5.4) and (5.5) give

$$r_{+} - t_{+} + 1 = \pi (i T_{11} - T_{12}) \tag{5.6}$$

and

$$r_{+} + t_{+} - 1 = -\pi (T_{21} + iT_{22}). \tag{5.7}$$

Similarly, for a regular wave from  $x = -\infty$ , we have

$$\phi_{\rm I} = A_{\rm -} \, \mathrm{e}^{-Ky + \mathrm{i} \, Kx} \quad \text{and} \quad \phi(P) \sim \begin{cases} A_{\rm -} \, t_{\rm -} \, \mathrm{e}^{-Ky + \mathrm{i} \, Kx} & \text{as } x \to +\infty, \\ A_{\rm -} \, \mathrm{e}^{-Ky} (\mathrm{e}^{\mathrm{i} \, Kx} + r_{\rm -} \mathrm{e}^{-\mathrm{i} \, Kx}) & \text{as } x \to -\infty. \end{cases}$$
(5.8)

In this case, we have

$$d_1 = -\mathrm{i}\pi A_-, \qquad d_2 = -\pi A_-$$

and  $d_n = 0$  for n > 2. It follows that

$$r_{-} - t_{-} + 1 = \pi (i T_{11} + T_{12}) \tag{5.9}$$

and

$$r_{-} + t_{-} - 1 = \pi (T_{21} - iT_{22}).$$
 (5.10)

If we eliminate  $r_+$  between (5.6) and (5.7), and  $r_-$  between (5.9) and (5.10), we obtain

$$t_{+}-t_{-}=\pi(T_{12}-T_{21})$$

whence the symmetry of the T-matrix, (4.1), shows that

$$t_{+} = t_{-} = t$$
, say. (5.11)

From (5.6), (5.7), (5.9) and (5.10), we have

$$\pi i T_{11} = 1 - t + \frac{1}{2}(r_- + r_+),$$
 (5.12)

$$\pi i T_{22} = 1 - t - \frac{1}{2}(r_- + r_+),$$
 (5.13)

and

$$\pi T_{12} = \pi T_{21} = \frac{1}{2}(r_{-} - r_{+}). \tag{5.14}$$

If we substitute these into (4.5), we obtain

$$|r_{-} + r_{+} - 2t|^{2} + |r_{-} - r_{+}|^{2} - 4 = 0,$$
 (5.15)

$$|r_{-} + r_{+} + 2t|^{2} + |r_{-} - r_{+}|^{2} - 4 = 0,$$
 (5.16)

and

$$|r_{-}|^{2} - |r_{+}|^{2} + 2i \operatorname{Im}(tr_{+}^{*} + t^{*}r_{-}) = 0,$$
 (5.17)

where (5.15), (5.16) and (5.17) correspond to m = n = 1, m = n = 2 and m = 1, n = 2 (or m = 2, n = 1), respectively. Taking the real and imaginary parts of (5.17), gives

$$|r_{+}| = |r_{-}| \tag{5.18}$$

and

$$Im(tr_{+}^{*} + tr_{-}^{*}) = 0. {(5.19)}$$

Adding (5.15) and (5.16) gives

$$|r_{+}|^{2} + |r_{-}|^{2} + 2|t|^{2} = 2$$

whence (5.18) implies that

$$|r_{+}|^{2} + |t|^{2} = 1 ag{5.20}_{+}$$

and

$$|r_-|^2 + |t|^2 = 1.$$
 (5.20\_)

Subtracting (5.15) and (5.16) gives

$$Re(tr_{+}^*+t^*r_{-})=0;$$

combining this with (5.19) shows that

$$tr_{+}^{*} + t^{*}r_{-} = 0. {(5.21)}$$

Using (5.18), (5.21) can be rewritten as

$$2 \arg(t) - \arg(r_+) - \arg(r_-) = \pi \mod 2\pi. \tag{5.22}$$

The results (5.11), (5.18), (5.20) and (5.22) are all well known: we shall call them the *Kreisel-Meyer relations*; they are derived systematically by Newman [10] for water of constant finite depth (he also derives the corresponding results in three dimensions). Kreisel [19] obtained (5.18) and (5.20), and their generalisations to the situation where the asymptotic depths of water are different at  $x = \pm \infty$ . R. Meyer, in an appendix to a paper by Biesel and Le Méhauté [20], proved (5.11), (5.18), (5.20) and (5.22). He also showed that, by a suitable shift of origin, it can be arranged that  $\arg(r_+) = \arg(r_-)$ , whence  $r_+ = r_-$ ; however, this shift cannot be determined a priori and, moreover, it depends on K.

# 6. Extension to finite depth, using multipole potentials

We can obtain similar results when the water is of constant finite depth, h, say. We begin by modifying S, so that the condition on  $\phi$  as  $y \to \infty$  is replaced by

$$\frac{\partial \phi}{\partial n} = 0$$
 on the bottom,  $y = h$ . (6.1)

The potential of a simple wave source is now given by [17]

$$G(x, y; \xi, \eta) = \frac{1}{2} \log \frac{(x - \xi)^2 + (y - \eta)^2}{(x - \xi)^2 + (y + \eta)^2}$$

$$-2 \oint_0^\infty \frac{\cosh k(h - y) \cosh k(h - \eta) \cos k(x - \xi) dk}{\cosh kh(k \sinh kh - K \cosh kh)}$$

$$-2 \int_0^\infty e^{-kh} \frac{\sinh ky \sinh k\eta}{k \cosh kh} \cos k(x - \xi) dk$$
(6.2)

where the path of integration passes below the pole of the integrand at  $k = k_0$ , and  $k_0$  is the unique positive real root of

$$K = k_0 \tanh k_0 h. \tag{6.3}$$

G has the bilinear expansion [17]

$$G(P,Q) = \alpha_m(P)\tilde{\Phi}_m(Q) \tag{6.4}$$

for  $r_P < r_Q$ , where the functions  $\alpha_m$  are the same as for the case of infinite depth, and  $\tilde{\Phi}_m$  are the finite-depth multipole potentials;  $\tilde{\Phi}_m$  are defined in Appendix A—they are harmonic, satisfy the free-surface and radiation conditions, satisfy the bottom condition (6.1), and are singular at O. Using (6.4), we obtain all the formulae derived in Section 3, with  $\Phi_m$  replaced by  $\tilde{\Phi}_m$  (except that we do not have to make this replacement in (3.9)). In particular, we have

$$2\pi\phi_{D}(P) = c_{m}\tilde{\Phi}_{m}(P), \quad P \text{ outside } C_{+}$$
(6.5)

and

$$2\pi\phi_{\mathbf{I}}(P_{-}) = d_{m}\alpha_{m}(P_{-}), \quad P \text{ inside } C_{-}$$

$$\tag{6.6}$$

with

$$c_m = T_{mn}d_n, \quad m = 1, 2, \dots \tag{6.7}$$

As before, the reciprocity relation implies that T is symmetric. The energy relation yields

$$T_{jn}[\alpha_m^*, \tilde{\Phi}_j] - T_{jm}^*[\alpha_n, \tilde{\Phi}_j^*] + T_{km}^* T_{jn}[\tilde{\Phi}_k^*, \tilde{\Phi}_j] = 0.$$
(6.8)

Ursell [17] has evaluated the bilinear products occurring here (see Appendix A); when these are substituted into (6.8), we eventually obtain

$$\frac{2}{\pi} \operatorname{Im}(T_{mn}) + A(P_m P_n^* + Q_m Q_n^*) = 0 \tag{6.9}$$

where

$$A(k_0h) = \frac{2\cosh^2 k_0h}{2k_0h + \sinh 2k_0h},\tag{6.10}$$

$$P_{m} = \frac{k_{0}}{K} T_{1m} + \operatorname{sech}^{2} k_{0} h \sum_{j=1}^{\infty} \frac{k_{0}^{2j+1}}{(2j)!} T_{m,2j+1}$$
(6.11)

and

$$Q_m = T_{2m} + \operatorname{sech}^2 k_0 h \sum_{j=1}^{\infty} \frac{k_0^{2j}}{(2j-1)!} T_{m,2j+2}.$$
 (6.12)

(6.9) is a finite-depth form of (4.5); moreover, as  $h \to \infty$ , (6.9) reduces to (4.5).

Consider a regular wave from  $x = +\infty$ , with potential

$$\phi_1 = A_+ Y_0(y) e^{-ik_0x}$$

where

$$Y_0(y) = \cosh k_0(h - y) \operatorname{sech} k_0 h,$$
 (6.13)

and define  $r_+$  and  $t_+$  by

$$\phi(P) \sim \begin{cases} A_{+} Y_{0}(y) (e^{-ik_{0}x} + r_{+} e^{ik_{0}x}) & \text{as } x \to +\infty, \\ A_{+} Y_{0}(y) t_{+} e^{-ik_{0}x} & \text{as } x \to -\infty. \end{cases}$$

it is shown in Appendix B that the coefficients  $d_m$  in (6.6) are given by

$$d_1 = i\pi(k_0/K)A_+, \qquad (2m)!d_{2m+1} = i\pi A_+ k_0^{2m+1} \operatorname{sech}^2 k_0 h,$$
  
$$d_2 = -\pi A_+, \qquad (2m-1)!d_{2m+2} = -\pi A_+ k_0^{2m} \operatorname{sech}^2 k_0 h,$$

for  $m \ge 1$ . Complicated expressions for  $r_+$  and  $t_+$  can then be found by using (6.5), (6.7) and the asymptotic behaviour of the multipole potentials,  $\tilde{\Phi}_m$  (see Appendix A); cf. the derivation of (5.6) and (5.7). Similarly, for a regular wave from  $x = -\infty$ ,

$$\phi_1 = A_- Y_0(y) e^{ik_0 x}$$

and  $r_{-}$  and  $t_{-}$  are defined by

$$\phi(P) \sim \begin{cases} A_{-}Y_{0}(y)t_{-} e^{ik_{0}x} & \text{as } x \to +\infty, \\ A_{-}Y_{0}(y)(e^{ik_{0}x} + r_{-} e^{-ik_{0}x}) & \text{as } x \to -\infty. \end{cases}$$

The corresponding coefficients  $d_m$  are given by (see Appendix B)

$$d_1 = -i\pi(k_0/K)A_-, \qquad (2m)!d_{2m+1} = -i\pi A_- k_0^{2m+1} \operatorname{sech}^2 k_0 h,$$
  

$$d_2 = -\pi A_-, \qquad (2m-1)!d_{2m+2} = -\pi A_- k_0^{2m} \operatorname{sech}^2 k_0 h,$$

for  $m \ge 1$ , and expressions for  $r_-$  and  $t_-$  can be found. The symmetry of T then implies that  $t_+ = t_- = t$ , say. We then obtain

$$1 - t + \frac{1}{2}(r_{-} + r_{+}) = A\pi i \left\{ \frac{k_0}{K} P_1 + \operatorname{sech}^2 k_0 h \sum_{n=1}^{\infty} \frac{k_0^{2n+1}}{(2n)!} P_{2n+1} \right\},$$
 (6.14)

$$1 - t - \frac{1}{2}(r_{-} + r_{+}) = A\pi i \left\{ Q_{2} + \operatorname{sech}^{2} k_{0} h \sum_{n=1}^{\infty} \frac{k_{0}^{2n}}{(2n-1)!} Q_{2n+2} \right\},$$
 (6.15)

and two equivalent expressions for  $r_- - r_+$ :

$$\frac{1}{2}(r_{-}-r_{+}) = A\pi \left\{ P_{2} + \operatorname{sech}^{2} k_{0} h \sum_{n=1}^{\infty} \frac{k_{0}^{2n}}{(2n-1)!} P_{2n+2} \right\}$$
(6.16)

$$= A\pi \left\{ \frac{k_0}{K} Q_1 + \operatorname{sech}^2 k_0 h \sum_{n=1}^{\infty} \frac{k_0^{2n+1}}{(2n)!} Q_{2n+1} \right\}.$$
 (6.17)

We can use these, together with (6.9), to verify that  $r_+$ ,  $r_-$  and t satisfy the Kreisel-Meyer relations. For example, consider (5.15): we have

$$|r_{-} + r_{+} - 2t|^{2} + |r_{-} - r_{+}|^{2} - 4 = |\lambda|^{2} + |r_{-} - r_{+}|^{2} - 2(\lambda + \lambda^{*})$$

$$(6.18)$$

where

$$\lambda = 2 - 2t + r_{-} + r_{+}$$

From (6.11) and (6.14), we have

$$2(\lambda + \lambda^*) = -8A\pi \left\{ \frac{k_0^2}{K^2} \operatorname{Im}(T_{11}) + \frac{2k_0}{K} \operatorname{sech}^2 k_0 h \sum_{j=1}^{\infty} \frac{k_0^{2j+1}}{(2j)!} \operatorname{Im}(T_{1,2j+1}) + \operatorname{sech}^4 k_0 h \sum_{j=1}^{\infty} \sum_{k=1}^{\infty} \frac{k_0^{2j+2k+2}}{(2j)!(2k)!} \operatorname{Im}(T_{2k+1,2j+1}) \right\}.$$

Using (6.9), we obtain

$$2(\lambda + \lambda^*) = (2A\pi)^2 \left\{ \frac{k_0^2}{k^2} |P_1|^2 + \frac{k_0}{K} \operatorname{sech}^2 k_0 h \sum_{j=1}^{\infty} \frac{k_0^{2j+1}}{(2j)!} (P_1 P_{2j+1}^* + P_{2j+1} P_1^*) + \operatorname{sech}^4 k_0 h \sum_{j=1}^{\infty} \sum_{k=1}^{\infty} \frac{k_0^{2j+2k+2}}{(2j)!(2k)!} P_{2k+1} P_{2j+1}^* + \operatorname{similar terms with } P_n \text{ replaced by } Q_n \right\}.$$

When this is substituted into (6.18), together with (6.14) and (6.17), it is seen that (5.15) is satisfied.

### 7. Extension to finite depth, using Havelock wavemaker functions

Most of the formulae in Section 6 for water of finite depth are very complicated, whereas the corresponding formulae for infinite depth are simple, e.g., compare (4.5) with (6.9). The reason for this is that all of the multipole potentials  $\tilde{\Phi}_m$  are wave-like at infinity, whereas only  $\Phi_1$  and  $\Phi_2$  generate waves (see Appendix A). Thus, we wish to replace  $\{\tilde{\Phi}_m\}$  with a different set of functions, most of which are wavefree (i.e. evanescent). The relevant functions are the Havelock wavemaker functions; the functions corresponding to  $\{\alpha_m\}$  are suggested by the following alternative bilinear expansion for G (defined by (6.2)), which is given by, e.g. John [2],

$$G(P,Q) = \beta_m(P)\Psi_m(Q) \tag{7.1}$$

for  $|x| < |\xi|$ , where

$$\begin{split} \Psi_1(P) &= A \pi \operatorname{sgn}(x) \, Y_0(y) \, \mathrm{e}^{\mathrm{i} k_0 |x|}, \qquad \Psi_2(P) = A \pi \mathrm{i} \, Y_0(y) \, \mathrm{e}^{\mathrm{i} k_0 |x|}, \\ \Psi_{2m+1}(P) &= A_m \pi \operatorname{sgn}(x) \, Y_m(y) \, \mathrm{e}^{-k_m |x|}, \qquad \Psi_{2m+2}(P) = A_m \pi \, Y_m(y) \, \mathrm{e}^{-k_m |x|}, \\ \beta_1(P) &= -2 \, Y_0(y) \, \sin k_0 x, \qquad \beta_2(P) = -2 \, Y_0(y) \, \cos k_0 x, \\ \beta_{2m+1}(P) &= -2 \, Y_m(y) \, \sinh k_m x, \qquad \beta_{2m+2}(P) = -2 \, Y_m(y) \, \cosh k_m x. \end{split}$$

Here, A and  $Y_0$  are defined by (6.10) and (6.13), respectively,  $k_m$ , m = 1, 2, ..., are the positive real solutions of

$$K + k_m \tan k_m h = 0$$
,  $A_m = \frac{2\cos^2 k_m h}{2k_m h + \sin 2k_m h}$  and  $Y_m(\hat{y}) = \frac{\cos k_m (h - y)}{\cos k_m h}$ .

 $\Psi_m$  are (linear combinations of) the well-known Havelock wavemaker functions [21, 22]: consider a semi-infinite channel x > 0 with a wavemaker at x = 0; suppose that  $\partial \phi / \partial x = f(y)$  on x = 0, 0 < y < h; then

$$\phi(x, y) = \sum_{m=0}^{\infty} a_m \Psi_{2m+2}(x, y)$$
 where  $a_m = -\frac{2}{\pi} \int_0^h f(y) Y_m(y) dy$ ,

since

$$\int_{0}^{h} Y_{m}(y) Y_{n}(y) dy = \frac{\delta_{nm}}{2k_{m}A_{m}}$$
 (7.2)

for m,  $n \ge 0$ , with  $A_0 = A$ .  $\Psi_m$  and  $\beta_m$  are harmonic functions that also satisfy the free-surface and bottom conditions;  $\beta_m$  are regular but  $\Psi_m$  are not smooth at x = 0;  $\Psi_m$  satisfy the radiation conditions as  $x \to \pm \infty$ . Suppose that

$$2\pi\phi_{\mathcal{D}}(P) = \tilde{c}_m \Psi_m(P) \quad \text{for } |x| > X \tag{7.3}$$

and

$$2\pi\phi_{\mathrm{I}}(P) = \tilde{d}_{m}\beta_{m}(P) \quad \text{for } |x| < X, \tag{7.4}$$

where X is a positive constant. Let

$$\tilde{c}_m = T_{mn}\tilde{d}_n, \quad m = 1, 2, \dots \tag{7.5}$$

for some  $T_{mn}$ . Reciprocity then implies that T is symmetric, as before, whilst it can be shown that the energy relation (2.5) yields

$$\frac{2}{\pi} \operatorname{Im}(T_{mn}) + A(T_{1m}T_{1n}^* + T_{2m}T_{2n}^*) = 0, \tag{7.6}$$

which should be compared with (4.5).

For |x| > X, (3.2) and (7.1) give (7.3) with

$$\tilde{c}_m = -\langle \phi, \beta_m \rangle. \tag{7.7}$$

However, for a *floating* body, we cannot use the same procedure to reduce (3.3) to (7.4) (since there will always be points  $q = (\xi, \eta) \in \partial D$  with  $|\xi| < |x|$ ). Instead, suppose we use (6.4) to give (6.6), namely

$$2\pi\phi_{\mathbf{I}}(P_{-}) = d_{m}\alpha_{m}(P_{-}),\tag{7.8}$$

with

$$d_m = \langle \phi, \tilde{\Phi}_m \rangle, \quad m = 1, 2, \dots; \tag{7.9}$$

these are the null-field equations for finite depth. Using (3.10) and eliminating  $a_n$ ,  $n = 1, 2, \dots$ , we obtain

$$\tilde{c}_m = \tilde{T}_{mn} d_n \tag{7.10}$$

where

$$\tilde{T}_{mn} = -\hat{R}_{mk}\tilde{Q}_{kn}^{-1}, \qquad \tilde{Q}_{mn} = \langle \phi_n, \tilde{\Phi}_m \rangle \quad \text{and} \quad \hat{R}_{mn} = \langle \phi_n, \beta_m \rangle.$$

Thus, we have a viable procedure for computing (an approximation to)  $\tilde{T}_{mn}$ . Now, if we can relate (7.4) and (7.8), i.e. if we can find the matrix S where

$$\tilde{d}_m = S_{mn}d_n, \quad m = 1, 2, \ldots,$$

then, comparing (7.5) and (7.10), we obtain

$$\tilde{T}_{mn} = T_{mk}S_{kn}, \quad m, n = 1, 2, \dots$$
 (7.11)

Properties of  $\tilde{T}$  can then be derived from those for T.

The matrix S is known explicitly: suppose

$$\tilde{d}_m \beta_m(x, y) = d_m \alpha_m(x, y); \tag{7.12}$$

set x = 0 to give

$$-2\sum_{m=0}^{\infty}\tilde{d}_{2m+2}Y_m(y)=\sum_{n=1}^{\infty}d_{2n}\alpha_{2n}(0,y);$$

the orthogonality of  $\{Y_m(y)\}$ , (7.2), then yields

$$\tilde{d}_{2m+2} = -k_m A_m \sum_{n=1}^{\infty} d_{2n} \int_0^h \alpha_{2n}(0, y) Y_m(y) dy, \quad m = 0, 1, 2, \dots;$$

 $\tilde{d}_{2m+1}$  are obtained by first differentiating (7.12) with respect to x. Actually, it is also possible to calculate  $S^{-1}$  explicitly; for a typical example, see Appendix B.

Finally, we observe that if  $\partial D$  intersects the bottom (physically, this means that the fluid domain is split into two separate domains, each one of which corresponds to a 'generalised wavemaker problem', i.e. a semi-infinite channel with an end-wall or arbitrary shape), then

$$T_{mn} = -\hat{R}_{mk}R_{kn}^{-1}$$
 where  $R_{mn} = \langle \phi_n, \Psi_m \rangle$ .

The corresponding null-field equations are

$$\tilde{d}_m = \langle \phi, \Psi_m \rangle, \quad m = 1, 2, \dots, \tag{7.13}$$

and these could be used to solve the generalised wavemaker problem. Furthermore, if  $\phi_1$  corresponds to a regular surface wave, the method of projection [4] can be used to yield a convergent numerical scheme for solving (7.13). This approach should also be useful for treating various three-dimensional problems, e.g., scattering by a right circular cone which is resting on the bottom and which pierces the free surface. These new methods are currently under investigation.

#### Appendix A. Multipole potentials

From [17], we have (3.4) with

$$\Phi_{2}(P) = \oint_{0}^{\infty} e^{-ky} \cos kx \frac{dk}{k - K}, \qquad \Phi_{1}(P) = -\frac{1}{K} \frac{\partial}{\partial x} \Phi_{2}(P),$$

$$\Phi_{2m+2}(P) = \frac{\cos 2m\theta}{r^{2m}} + \frac{K}{2m - 1} \frac{\cos(2m - 1)\theta}{r^{2m - 1}}, \qquad \Phi_{2m+1}(P) = \frac{\sin(2m + 1)\theta}{r^{2m + 1}} + \frac{K}{2m} \frac{\sin 2m\theta}{r^{2m}},$$

$$\alpha_{2}(P) = -2 e^{-Ky} \cos Kx, \qquad \alpha_{1}(P) = -2 e^{-Ky} \sin Kx,$$

$$\alpha_{2m+2}(P) = \frac{-2(2m - 1)!}{K^{2m}} \sum_{q=2m}^{\infty} \frac{(-Kr)^{q}}{q!} \cos q\theta, \qquad \alpha_{2m+1}(P) = \frac{2(2m)!}{K^{2m+1}} \sum_{q=2m+1}^{\infty} \frac{(-Kr)^{q}}{q!} \sin q\theta,$$
(A.1)

 $m = 1, 2, \ldots$ , and the point P = (x, y) has circular polar coordinates given by  $x = r \sin \theta$ ,  $y = r \cos \theta$  (with  $r = r_p$ ). Note that  $\Phi_{2m}$  and  $\alpha_{2m}$  ( $\Phi_{2m-1}$  and  $\alpha_{2m-1}$ ) are even (odd) functions of x,  $m = 1, 2, \ldots$ . For water of constant finite depth, h, say, the multipole potentials are given by [17]

$$\begin{split} \tilde{\Phi}_{2}(P) &= \int_{0}^{\infty} \frac{\cosh k(h-y) \cos kx \, \mathrm{d}k}{k \sinh kh - K \cosh kh}, \qquad \tilde{\Phi}_{1}(P) = -\frac{1}{K} \frac{\partial}{\partial x} \, \tilde{\Phi}_{2}(P), \\ \tilde{\Phi}_{2m+2}(P) &= \Phi_{2m+2}(P) - \frac{1}{(2m-1)!} \int_{0}^{\infty} \frac{\mathrm{e}^{-kh}(K+k)(K \sinh ky - k \cosh ky)k^{2m-2} \cos kx \, \mathrm{d}k}{k \sinh kh - K \cosh kh}, \\ \tilde{\Phi}_{2m+1}(P) &= \Phi_{2m+1}(P) - \frac{1}{(2m)!} \int_{0}^{\infty} \frac{\mathrm{e}^{-kh}(K+k)(K \sinh ky - k \cosh ky)k^{2m-1} \sin kx \, \mathrm{d}k}{k \sinh kh - K \cosh kh}. \end{split}$$

Asymptotic behaviour of the multipole potentials as  $|x| \to \infty$ 

We have

$$\Phi_1 \sim \pm \pi e^{-Ky \pm iKx}$$
 and  $\Phi_2 \sim \pi i e^{-Ky \pm iKx}$  as  $x \to \pm \infty$ .

For m > 2,  $\Phi_m$  is a wavefree potential, i.e. it decays algebraically as  $|x| \to \infty$ .

For finite depth, we have

$$\tilde{\Phi}_{1} \sim \pm \pi (k_{0}/K) A Y_{0} e^{\pm ik_{0}x}, \qquad (2m)! \tilde{\Phi}_{2m+1} \sim \pm \pi k_{0}^{2m+1} \operatorname{sech}^{2}(k_{0}h) A Y_{0} e^{\pm ik_{0}x}, 
\tilde{\Phi}_{2} \sim \pi i A Y_{0} e^{\pm ik_{0}x}, \qquad (2m-1)! \tilde{\Phi}_{2m+2} \sim \pi i k_{0}^{2m} \operatorname{sech}^{2}(k_{0}h) A Y_{0} e^{\pm ik_{0}x}$$

as  $x \to \pm \infty$ , where  $k_0$ , A and  $Y_0$  are defined by (6.3), (6.10) and (6.13), respectively, and  $m = 1, 2, \ldots$ (Note that, as  $h \to \infty$ ,  $k_0 \to K$ ,  $A \to 1$  and, for fixed  $y \ (0 \le y < h)$ ,  $Y_0 \to e^{-Ky}$ .)

### Bilinear products

From [17], we have

$$[\alpha_m^*, \tilde{\Phi}_j] = [\alpha_m, \tilde{\Phi}_j^*] = 2\pi\delta_{jm}$$
 and  $[\tilde{\Phi}_k^*, \tilde{\Phi}_j] = 2\pi^2 i A \Omega_{kj}$ 

where  $\Omega_{kj} = \Omega_{jk}$ ,  $\Omega_{2m,2n+1} = 0$ ,

$$\Omega_{11} = k_0^2 / K^2, \qquad K(2m)! \Omega_{1,2m+1} = k_0^{2m+2} \operatorname{sech}^2 k_0 h,$$

$$\Omega_{22} = 1, \qquad (2m-1)! \Omega_{2,2m+2} = k_0^{2m} \operatorname{sech}^2 k_0 h$$

$$(2m)! (2n)! \Omega_{2m+1,2n+1} = k_0^{2m+2n+2} \operatorname{sech}^4 k_0 h$$

and

$$(2m-1)!(2n-1)!\Omega_{2m+2,2n+2} = k_0^{2m+2n} \operatorname{sech}^4 k_0 h.$$

# Appendix B

Consider a regular surface wave, propagating from  $x = +\infty$  on water of depth h. We seek the coefficients  $d_m^+$  in the expansion

$$2\pi\phi_{I}(x,y) = 2\pi Y_{0}(y) e^{-ik_{0}x} = d_{m}^{+}\alpha_{m}(x,y).$$
(B.1)

Set x = 0, whence

$$2\pi Y_0(y) = \sum_{n=0}^{\infty} d_{2n+2}^{+} \alpha_{2n+2}(0, y).$$
 (B.2)

We have

$$Y_0(y) = \cosh k_0 y - (K/k_0) \sinh k_0 y = 1 - Ky + \sum_{m=1}^{\infty} p_m(Ky)(k_0/K)^{2m}$$
(B.3)

where

$$p_m(Ky) = \frac{(Ky)^{2m}}{(2m)!} \left(1 - \frac{Ky}{2m+1}\right).$$

From Appendix A, we have

$$\alpha_2(0, y) = -2\left(1 - Ky + \sum_{m=1}^{\infty} p_m(Ky)\right)$$
(B.4)

and

$$\alpha_{2n+2}(0, y) = -\frac{2(2n-1)!}{K^{2n}} \sum_{m=n}^{\infty} p_m(Ky)$$

for  $n = 1, 2, \ldots$  Hence

$$\sum_{n=1}^{\infty} d_{2n+2}^{+} \alpha_{2n+2}(0, y) = -2 \sum_{m=1}^{\infty} p_m(Ky) \sum_{n=1}^{m} \frac{(2n-1)!}{K^{2n}} d_{2n+2}^{+}.$$
 (B.5)

Substituting (B.3)-(B.5) into (B.2), and comparing coefficients gives

$$2\pi = -2d_2^+$$
 and  $2\pi \left(\frac{k_0}{K}\right)^{2m} = -2\left\{d_2^+ + \sum_{n=1}^m \frac{(2n-1)!}{K^{2n}} d_{2n+2}^+\right\}$ 

for  $m = 1, 2, \ldots$ . Hence  $d_2^+ = -\pi$  and

$$1 - \left(\frac{k_0}{K}\right)^{2m} = \frac{1}{\pi} \sum_{n=1}^{m} \frac{(2n-1)!}{K^{2n}} d_{2n+2}^{+} \quad \text{for } m = 1, 2, \dots$$

Since

$$1-x^{m}=(1-x)\sum_{n=1}^{m}x^{n-1},$$

.

we obtain

$$d_{2n+2}^{+} = \frac{\pi(K^2 - k_0^2)k_0^{2n-2}}{(2n-1)!} = -\frac{\pi k_0^{2n} \operatorname{sech}^2 k_0 h}{(2n-1)!}.$$

The odd coefficients can be obtained by differentiating (B.1) with respect to x noting that

$$\frac{\partial}{\partial x} \alpha_1 = K\alpha_2$$
 and  $\frac{\partial}{\partial x} \alpha_{2n+1} = 2n\alpha_{2n+2}$ 

for n = 1, 2, ... and then setting x = 0; the results are

$$d_1^+ = \pi i (k_0/K)$$
 and  $(2n)! d_{2n+1}^+ = \pi i k_0^{2n+1} \operatorname{sech}^2 k_0 h$ .

For a regular wave from  $x = -\infty$ , we have

$$2\pi\phi_{\rm I}(x,y) \equiv 2\pi Y_0(y) e^{ik_0x} = d_m^-\alpha_m(x,y);$$

it is easy to see that

$$d_{2n+1}^- = -d_{2n+1}^+$$
 and  $d_{2n+2}^- = d_{2n+2}^+$  for  $n = 0, 1, \dots$ 

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